

Context and Motivation

- Accurate determination of the physical conditions in comets can be inferred from the modelling of observational spectra
- The full exploitation of these spectra requires going beyond the local thermodynamic equilibrium
 - Radiative and collisional properties are needed
- CO, CO₂, and H₂O are the most abundant molecules in comets
- For comets at large heliocentric distances, production of CO significantly exceeds production of H₂O [1]
- Accurate rate coefficients for CO-CO collisional system are essential for proper astrophysical modelling of comets

Details of the calculations

- The 4D Potential Energy Surface (PES) calculated by Visser et al. [2] was used in our calculations
- Calculations were performed within the coupled states (CS) approximation using the MOLSCAT scattering code [3]
- Molecules were treated as distinguishable ones for the purpose of astrophysical applications
- Cross sections were calculated for transitions between rotational levels of CO molecules up to $j_1 = j_2 = 10$
- From calculated cross-sections, we provide rate coefficients for temperatures up to 150 K
- Statistical Adiabatic Channel Model (SACM) was investigated to extend calculations to higher rotational levels [4]

Propensity rules

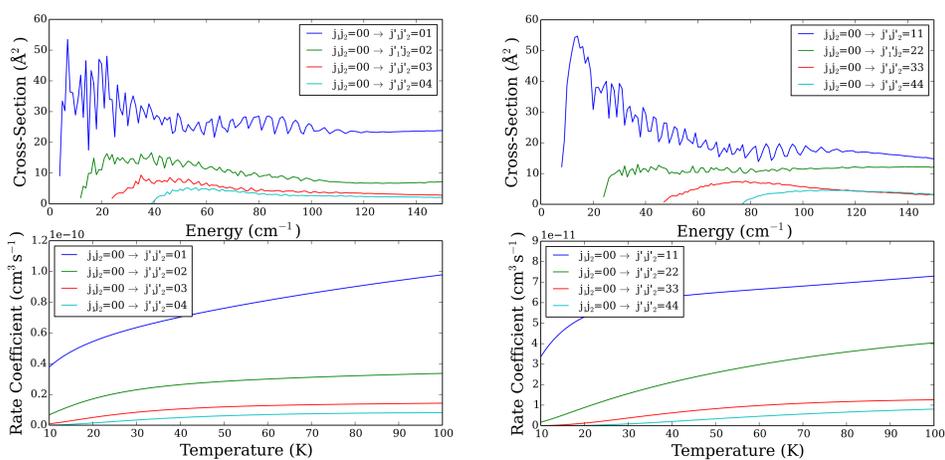


Fig.1 Cross-sections and rate coefficients of no-pair (only one CO is excited) transitions

Fig.2 Cross-sections and rate coefficients of pair (both CO are excited) transitions

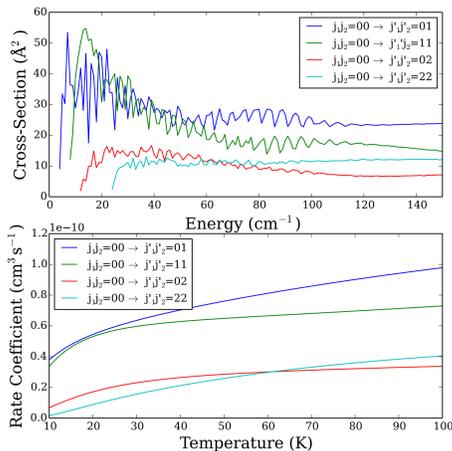


Fig.3 Comparison of cross-sections and rate coefficients between pair and no-pair transitions

- Strong propensity rules toward no pair transitions with $\Delta j_2 = 1$ (Fig.1)
- Cross sections decrease while Δj_2 increases
- Same behaviour for pair transitions (Fig.2)
- Similar magnitude of cross-sections and rate coefficients for pair and no pair transitions (Fig. 3)

Comparison with previous study

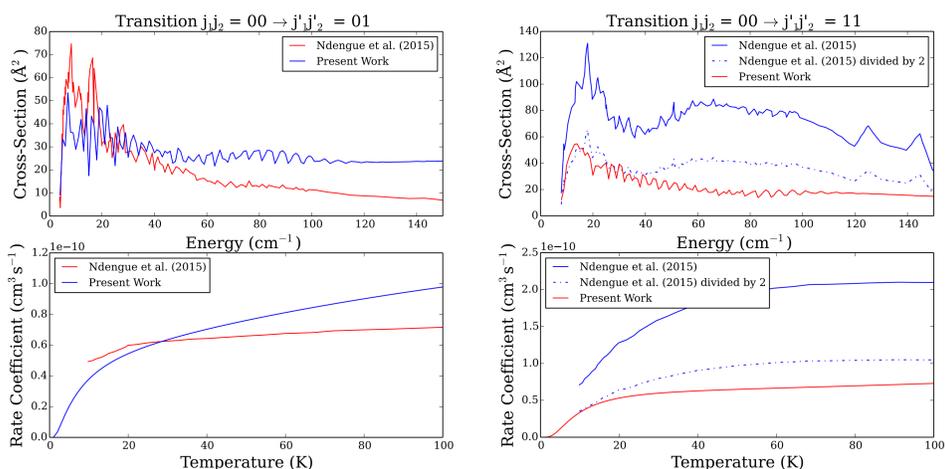


Fig.4 Comparison of cross-section and rate coefficients for first no-pair transitions

Fig.5 Comparison of cross-section and rate coefficients for first pair transitions

We observe differences due to:

- PESs used in calculations [2,6]
- Calculation method: close coupling (CC) + Multi Channel Time Depended Hartree (MCTDH) in previous work [5], coupled states approximation in this work
- Size of rotational basis: $j_1 = j_2 = 7$ in previous work [5], $j_1 = j_2 = 15$ in this work
- Error in the MOLSCAT code - overestimation of pair transitions (Fig 5.)

Statistical approach

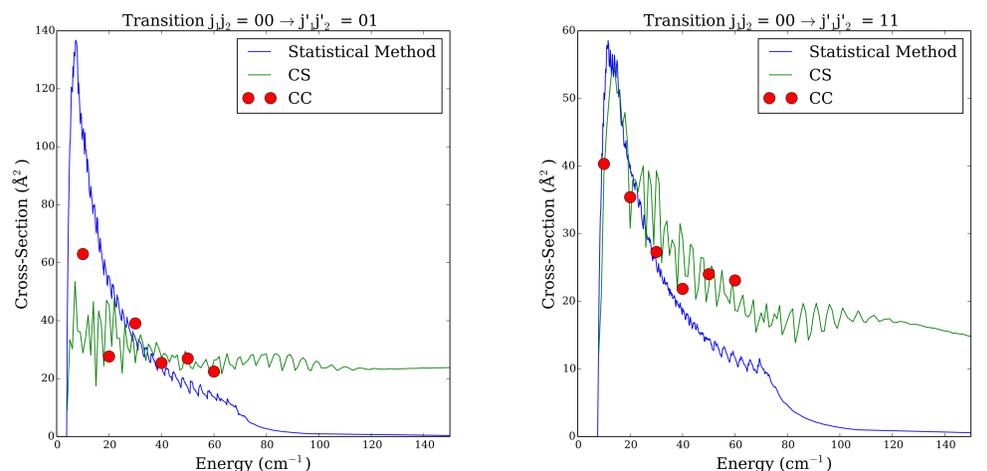


Fig.6 Comparison of cross-sections using different scattering methods

Fig.7 Comparison of cross-sections using different scattering methods

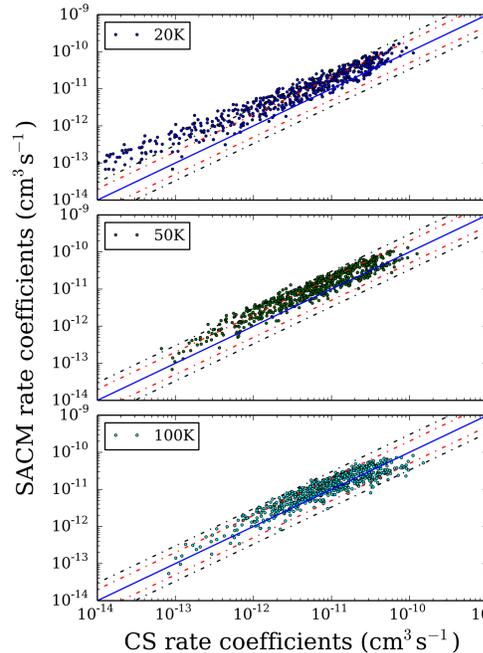


Fig.8 Systematic comparison of the rate coefficients for different temperatures. Blue line: identical results, red dashed line: factor of 2 difference, green dashed line: factor of 3 difference

- SCAM cross-sections are compared to the CS results
- Up to a factor of 3 of differences observed at low energy (Fig.6)
- Good agreement for part of the transitions at medium energy range (Fig. 7)
- Sudden decrease of cross-sections is observed above 60 cm⁻¹ (Fig. 6 and 7).
- Due to high density of accessible channels
- Systematic comparison of rate coefficients at 3 different temperatures (Fig. 8)
- 20K - Part of the transitions exceeds a factor of 3 differences
- 50K - Acceptable agreement within a factor of 3
- 100K - Good agreement but rather fortuitous at this temperature

Summary and Future plans

- We calculated rate coefficients for temperatures up to 150K for transitions between CO rotational levels $j_1 = j_2 = 10$
- Comparison with previous study uncover error in the MOLSCAT scattering code
 - The cross-sections are overestimated for pair transitions in collisional systems involving identical molecules
- SACM method does not work for the CO-CO system
 - Density of accessible channels is too high
- The new data will be used in the modelling of the CO rich comets

References

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- [4] J. Loreau et al., *ApJL* **853** L5 (2018)
- [5] S. Ndengué, R. Dawes, F. Gatti, *J. Chem. Phys.* **119**, (2015)
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