

Context

- Physical conditions in astrophysical media are derived from spectral line analysis
 - Requires the **population** of molecular energy levels
- Local thermodynamic equilibrium conditions rarely fulfilled
 - Need of **radiative** and **collisional** molecular data
- Rate coefficients, which characterize transitions induced by collisions, are obtained in 2 steps:
 - Interaction potential between the colliders
 - Potential energy surface (PES)
 - Scattering calculations based on the PES
 - Inelastic cross sections and **rate coefficients**

Potential energy surface

- UCCSD(T)/aVQZ level of theory using MOLPRO⁴
- Legendre polynomial expansion based on 1351 geometries

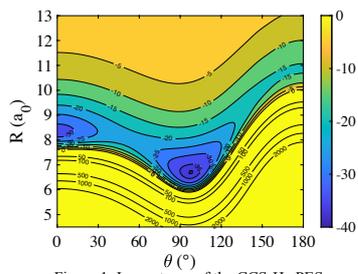


Figure 1: Isocontours of the CCS-He PES.

- Highly anisotropic
- Shallow minima ~ 35 cm⁻¹

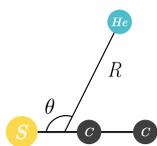


Figure 2: CCS-He collisional system in Jacobi coordinates.

CCS(Σ) fine structure

- $S = 1 \rightarrow j = |N - 1| ; N ; N + 1$

Electronic angular momentum S
Nuclear rotational momentum N
Total angular momentum $j = N + S$

1 rotational level N

→ 3 fine structure levels N_j

- Large spin-splitting
 - Energy levels **mixed up** up to $N_j = 10_j$

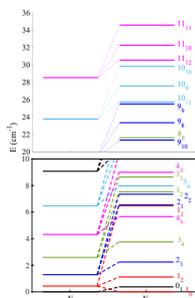


Figure 3: Energy levels of the CCS radical.

Rate coefficients

- Close-coupling approach using a modified version of MOLSCAT⁵
- Thermal averaging of the cross sections

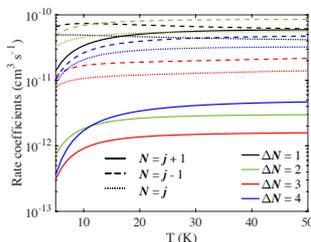


Figure 4: Fine structure resolved rate coefficients from $N_j = 1_j$ for $\Delta N = \Delta j$ transitions.

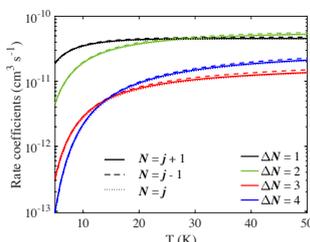


Figure 5: Fine structure resolved rate coefficients from $N_j = 10_j$ for $\Delta N = \Delta j$ transitions.

Fine structure:

- Must be **accurately** taken into account for **low N_j levels**
- Could be **approximatively** taken into account for **high N_j**

CCS - He collisional system

- CCS detected in several **molecular clouds** and in **IRC+10216**^{1,2}
- CC³⁴S**, **¹³CCS** and **C¹³CS** also detected in **several astronomical sources**
- He** is one of the **dominant collider** and a **proxy for H₂**

Objectives

- First PES for the CCS-He system
- First accurate **fine structure** resolved **rate coefficients** for the **CCS** and **CC³⁴S** isotopologues
- First **hyperfine structure** resolved **rate coefficients** for the **¹³CCS** and **C¹³CS** isotopologues

Effect of the isotopic substitution

Fine structure rate coefficients:

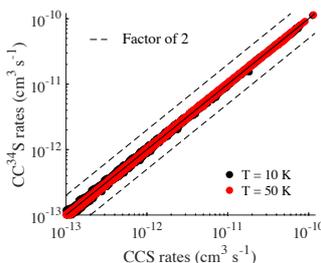


Figure 6: Systematic comparison of the rate coefficients of CCS and CC³⁴S isotopologues.

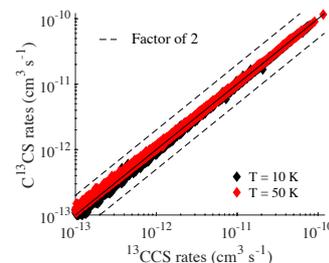


Figure 7: Systematic comparison of the rate coefficients of ¹³CCS and C¹³CS isotopologues.

→ Rate coefficients for **one isotopologues** can be used to infer the ones of all **4 isotopologues**

Hyperfine structure rate coefficients:

- $I = 1/2 \rightarrow F = |j - 1/2| ; j + 1/2$

Nuclear angular momentum I
Total angular momentum $F = j + I$

1 fine structure level N_j

→ 2 hyperfine structure levels N_j, F

- Recoupling method⁶

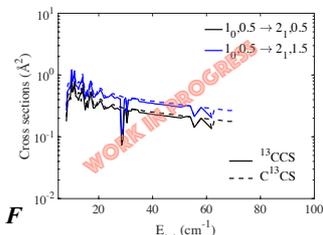


Figure 8: Hyperfine excitation cross sections of ¹³C-based isotopologues.

Conclusions and Perspectives

- ✓ CCS-He PES at the UCCSD(T)/aVQZ level of theory
- ✓ First **accurate** set of rate coefficients for the **5 - 50 K** temperature range
- ✓ **Isotopic substitution** has **no significant effect**

Hyperfine structure resolved rate coefficients for ¹³C-based isotopologues

- Modelisation of observation:
 - Observation of CCS
 - Determination of ¹³C-based isotopologues **abundances**

→ Investigation of the **chemical processes** leading to the **formation of CCS isotopologues**

References

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